

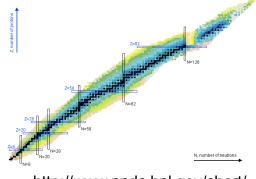
(NASA)

Nuclear Physics in One Lecture

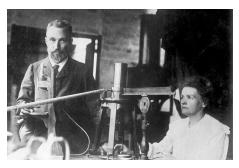
D. Schroeder, 26 November 2012

Outline

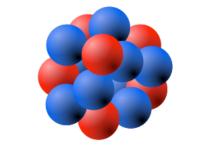
- History, clues, data
- Visualizing the nucleus
- Reasons for (in)stability
- Decay processes
- Fusion in the Sun



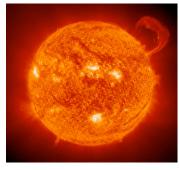
http://www.nndc.bnl.gov/chart/



(Wikimedia Commons)



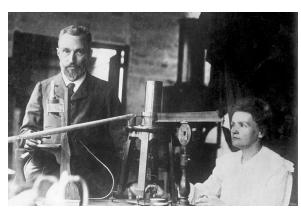
(Marekich, Wikimedia Commons)



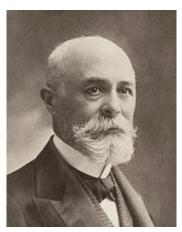
(NASA)

Discovery of Radioactivity

- Becquerel (1896): Uranium salts expose photographic plates
- Curies, Rutherford:
 - Better detectors
 - Transmutation of elements
 - $-\alpha$ vs. β



(Photo from Wikimedia Commons)



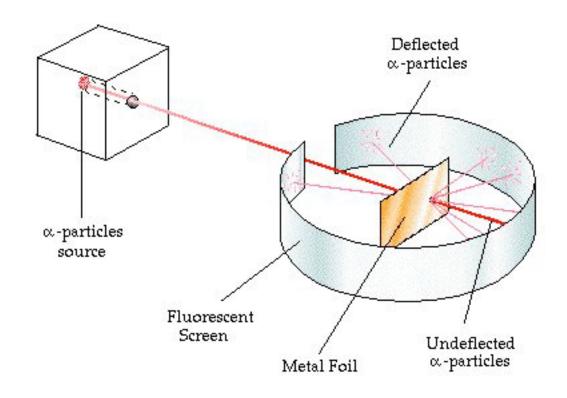
(Photo by Paul Nadar via Wikimedia Commons)



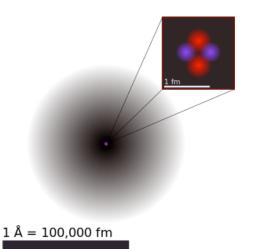
(Wikimedia Commons)

Discovery of the Nucleus

(Rutherford, Geiger, Marsden, 1911)



Scattering angles implied a tiny, massive center of force inside the atom.



(Drawing from www.wiredchemist.com)

(Yzmo, Wikimedia Commons)

What is the nucleus made of?

١.	luga in pariodia tabla:								2	He
, I l	lues in periodic table:								Helium	
								4.002602		
	6	С	7	N	8	О	9	F	10	Ne
	Carbon		Nitrogen		Oxygen		Fluorine		Neon	
	12.0107		14.00674		15.9994		18.9984032		20.1797	
1	14	Si	15	Р	16	S	17	CI	18	Ar
	Silicon		Phosph.		Sulfur		Chlorine		Argon	
	28.0855		30.973761		32.066		35.4527		39.948	
Ī	32 (Ge	33	As	34	Se	35	Br	36	Kr
	German. 72.61		Arsenic		Selenium		Bromine		Krypton	
			74.92160		78.96		79.904		83.80	
-	50	Sn	51	Sb	52	Te	53	I	54	Xe
	Tin 118.710		Antimony		Tellurium		Iodine		Xenon	
			121.760		127.60		126.90447		131.29	
1	82 F	Рb	83	Bi	84	Po	85	At	86	Rn
	Lead		Bismuth		Polonium		Astatine		Radon	
	207.2		208.98038		(208.982415)		(209.987131)		(222.017570)	

Masses in amu tend to be close to integers (A), about twice the atomic number (Z).

Rutherford discovered isolated protons coming from nuclei in 1917.

Chadwick discovered the neutron in 1932.

$$m_p \approx m_n \approx 1 \text{ u}$$

What is the nucleus made of?

1	lues in periodic table:								2	He	
, ,	idee in periodic table.									Helium	
										4.002602	
Ī	6	С	7	Ν	8	0	9	F	10	Ne	
	Carbon		Nitrogen		Oxygen		Fluorine		Neon		
	12.0107		14.00674		15.9994		18.9984032		20.1797		
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	Sil	Silicon 28.0855		Phosph.		Sulfur		Chlorine		Argon	
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	82	Pb	83	Bi	84	Po	85	At	86	Rn	
	Lead		Bismuth		Polonium		Astatine		Radon		
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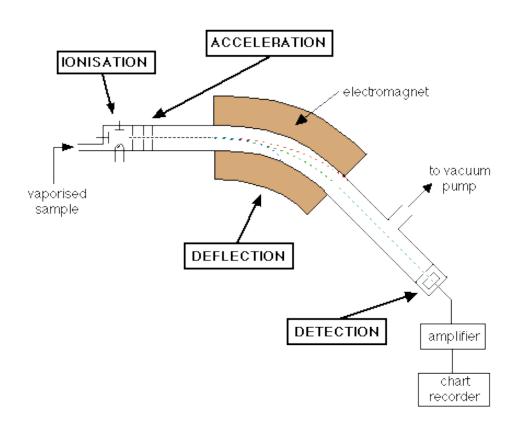
Non-integer masses are from mixtures of different *isotopes* (same number of protons, different numbers of neutrons).

Notation: ${}^{A}_{Z}X$

For example,

 ${}_{2}^{4}\mathrm{He}$ ${}_{6}^{14}\mathrm{C}$ ${}_{92}^{235}\mathrm{U}$

Mass Spectrometry



(Drawing from www.chemguide.co.uk)

Electrostatic acceleration:

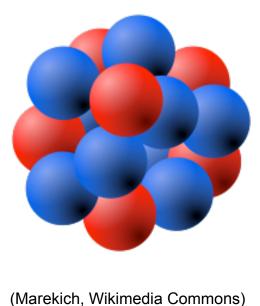
$$\frac{1}{2}mv^2 = q\,\Delta V$$

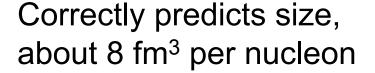
Magnetic bending:

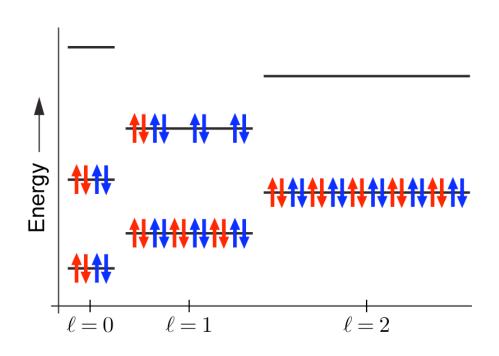
$$\frac{mv^2}{r} = qvB$$

Solve:
$$r=\sqrt{\frac{2m\,\Delta V}{qB^2}}$$

Two Wrong Pictures







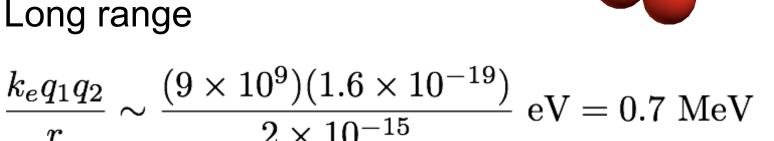
Fermions confined in a spherical potential well (many effects omitted!)

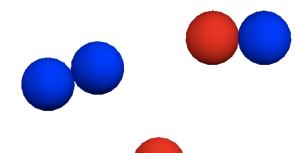
Reasons for (in)stability

- "Strong" force
 - Equally strong between pp, nn, pn
 - Short range



- Protons only
- Long range





Reasons for (in)stability

Quantum confinement energy

$$\frac{p^2}{2m} = \frac{h^2}{2m_p\lambda^2} \sim 100 \text{ MeV}$$

- Magnetic forces
- *p-n* conversion ("weak" interaction)

$$n \longrightarrow p + e^- + \bar{\nu}$$
 $p \longrightarrow n + e^+ + \nu$
 $(m_n - m_p = 1.3 \text{ MeV})$

Decay Processes

•
$$\alpha$$
 (⁴He)

$$^{238}_{92}U \longrightarrow ^{234}_{90}Th + ^{4}_{2}He$$

•
$$\beta^{-}$$
 (e⁻)

$$^{14}_{6}C \longrightarrow ^{14}_{7}N + ^{0}_{-1}e^{-} + ^{0}_{0}\bar{\nu}$$

•
$$\beta^{+}$$
 (e⁺)

$${}^{11}_{6}\mathrm{C} \longrightarrow {}^{11}_{5}\mathrm{B} + {}^{0}_{1}e^{+} + {}^{0}_{0}\nu$$

$$^{40}_{19}{
m K} + ^{0}_{-1}e^{-} \longrightarrow ^{40}_{18}{
m Ar} + ^{0}_{0}\nu$$

•
$$\gamma$$
 (photon)

$$^{99\text{m}}_{43}\text{Tc} \longrightarrow ^{99}_{43}\text{Tc} + \gamma$$

- All reactions conserve: electric charge, total number of nucleons.
- For spontaneous decays, total mass must decrease.

Decay Rates

 Each decay is a quantum process with a fixed probability per unit time, λ.

$$\frac{dN}{dt} = -\lambda N \quad \Rightarrow \quad N(t) = N(0) \cdot e^{-\lambda t}$$

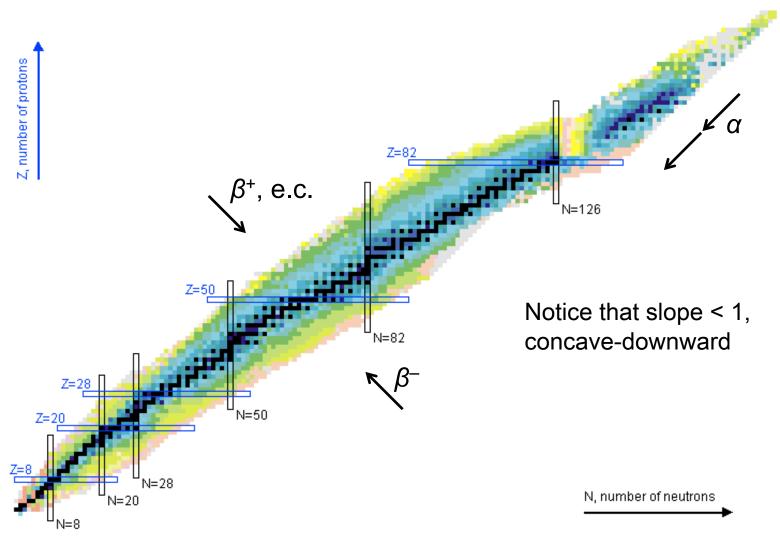
Half-life:

$$t_{1/2} = rac{\ln 2}{\lambda}$$



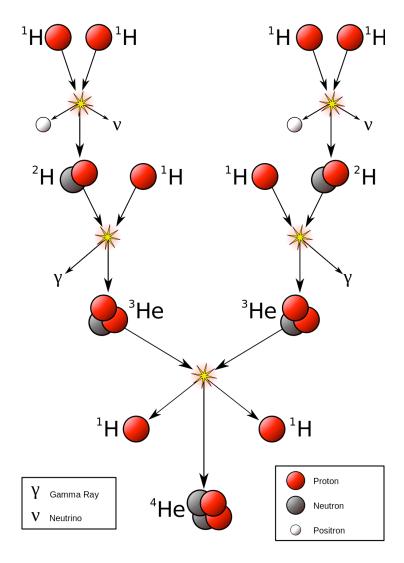
• $t_{1/2}$ varies from <10⁻¹⁵ s to >10²⁴ years!

Chart of the Nuclides



http://www.nndc.bnl.gov/chart/

Application: Fusion in the Sun



Net reaction:

$$4p + 2e^{-} \longrightarrow$$

$$^{4}\text{He} + 2\nu + 6\gamma$$

Mass loss: 0.7%

$$(\Delta m)c^2 = 26.7 \text{ MeV}$$

x 10^{38} reactions/second ≈ 4 x 10^{26} watts

(Borb, Wikimedia Commons)