

Homework #13(41-1) There are $2n^2$ electron states in shell n

a) $2n^2 = 2(4)^2 = 32$

b) $2n^2 = 2(1)^2 = 2$

c) $2n^2 = 2(3)^2 = 18$

d) $2n^2 = 2(2)^2 = 8$

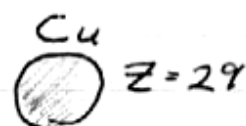
(41-2) $n = 1, l = 0, m_l = 0, m_s = +\frac{1}{2}$

$n = 1, l = 0, m_l = 0, m_s = -\frac{1}{2}$

(43-3) initial:  $\rightarrow$

$K_\alpha = 5.30 \text{ MeV}$

$U = 0$

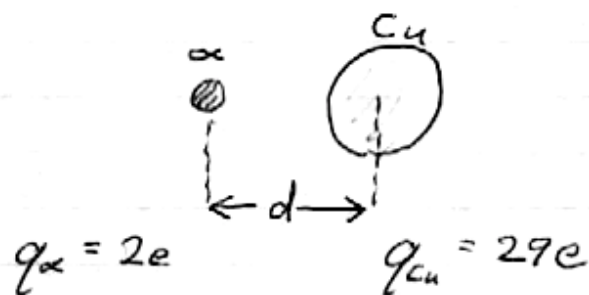


$K_{\text{Cu}} = 0$

final: $K_\alpha = 0$

$K_{\text{Cu}} = 0$

$U = \frac{1}{4\pi\epsilon_0} \frac{q_\alpha q_{\text{Cu}}}{d}$



By conservation of energy

$$5.30 \text{ MeV} = \frac{1}{4\pi\epsilon_0} \frac{q_\alpha q_{\text{Cu}}}{d}$$

$$d = \frac{1}{4\pi\epsilon_0} \frac{q_\alpha q_{\text{Cu}}}{5.30 \text{ MeV}}$$

Recalling that $\frac{1}{4\pi\epsilon_0} = 9 \times 10^9 \frac{\text{Nm}^2}{\text{C}^2}$, we find

$$d = \left(9 \times 10^9 \frac{\text{Nm}^2}{\text{C}^2} \right) \frac{(2 \times 29 \times 1.6 \times 10^{-19} \text{C})^2}{(5.30 \times 10^6 \text{eV} \times \frac{1.6 \times 10^{-19} \text{J}}{1 \text{eV}})}$$

$$= 1.58 \times 10^{-14} \text{m} = \boxed{15.8 \text{fm}}$$

(43-4) $\Delta x \Delta p \approx \hbar$ and $\Delta p = m \Delta v$ ($m = \text{proton or neutron mass}$)

So the uncertainty in the nucleon's velocity is

$$\Delta v \approx \frac{\hbar}{m \Delta x} = \frac{h/2\pi}{m \Delta x}$$

Using $\Delta x \approx R = R_0 A^{1/3} = 1.2 \text{fm} (100)^{1/3} = 5.57 \text{fm}$,
This is

$$\Delta v \approx \frac{(6.63 \times 10^{-34} \text{J} \cdot \text{sec})/2\pi}{(1.67 \times 10^{-27} \text{kg} \times 5.57 \times 10^{-15} \text{m})}$$

$$= 1.13 \times 10^7 \frac{\text{m}}{\text{sec}}$$

The value of the nucleon's speed must be at least $1.13 \times 10^7 \frac{\text{m}}{\text{sec}}$. Otherwise, since the uncertainty in the speed is $\pm 1.13 \times 10^7 \frac{\text{m}}{\text{sec}}$, we could have a negative value for a measurement of the speed. Setting $v = 1.13 \times 10^7 \frac{\text{m}}{\text{sec}}$, the kinetic energy is

$$K \approx \frac{1}{2} m v^2 = \frac{1}{2} (1.67 \times 10^{-27} \text{kg}) (1.13 \times 10^7 \frac{\text{m}}{\text{sec}})^2$$

or

$$K = 1.07 \times 10^{-13} \text{J} \left(\frac{1 \text{eV}}{1.6 \times 10^{-19} \text{J}} \right)$$

$$= \boxed{672 \text{keV}}$$

(43-5) The mass of a gallium atom, ^{67}Ga , is

$$m = 67u = 67(1.66 \times 10^{-27} \text{ kg}) = 1.11 \times 10^{-25} \text{ kg}$$

In this 3.4 g sample, the initial number of ^{67}Ga atoms is

$$N_0 = \frac{3.4 \times 10^{-3} \text{ kg}}{1.11 \times 10^{-25} \text{ kg/atom}} = 3.06 \times 10^{22} \text{ atoms}$$

The disintegration constant λ is

$$\lambda = \frac{\ln 2}{\tau} = \frac{\ln 2}{(78 \text{ hr})(3600 \frac{\text{sec}}{\text{hr}})} = 2.47 \times 10^{-6} \frac{1}{\text{sec}}$$

a) The initial decay rate is

$$R_0 = \lambda N_0 = (2.47 \times 10^{-6} \frac{1}{\text{sec}})(3.06 \times 10^{22} \text{ atoms})$$

$$\boxed{R_0 = 7.56 \times 10^{16} \frac{\text{decays}}{\text{sec}}}$$

$$b) R = R_0 e^{-\lambda t} = R_0 e^{-\ln(2)t/\tau}$$

$$\text{So } R = 7.56 \times 10^{16} \frac{\text{decays}}{\text{sec}} e^{-\ln(2)(48 \text{ hr})(78 \text{ hr})}$$

$$\boxed{R = 4.93 \times 10^{16} \frac{\text{decays}}{\text{sec}}}$$

$$(43-6) a) \tau = \frac{\ln 2}{\lambda} = \frac{\ln 2}{0.0108 \text{ hr}^{-1}} = \boxed{64.2 \text{ hr}}$$

b) After 3 half-lives, the number of ^{197}Hg atoms remaining is

$$N = \left(\frac{1}{2}\right)^3 N_0 = \frac{1}{8} N_0$$

One-eighth ~~of~~ of the sample remains.

$$c) \frac{N}{N_0} = e^{-\lambda t} = e^{-(0.0108 \text{ hr}^{-1} \times 10 \text{ d} \times \frac{24 \text{ hr}}{1 \text{ d}})} \\ = \boxed{0.0749}$$

$$(43-7) \quad Q = [(mass \text{ in}) - (mass \text{ out})] c^2 \\ = [m(^{238}\text{U}) - m(^{234}\text{Th}) - m(^4\text{He})] c^2 \\ = [238.05079 \text{ u} - 234.04363 \text{ u} - 4.00260 \text{ u}] c^2 \\ = 4.56 \times 10^{-3} \text{ u} c^2$$

Using $\text{u} c^2 = 931.5 \text{ MeV}$, we find

$$Q = (4.56 \times 10^{-3})(931.5 \text{ MeV}) \\ = \boxed{4.25 \text{ MeV}}$$

(44-8) a) The mass of the ^{235}U atom is

$$m = 235.04392 \text{ u} \\ = 235.04392 (1.661 \times 10^{-27} \text{ kg}) \\ = 3.904 \times 10^{-25} \text{ kg}$$

The number of ^{235}U atoms in 1 kg is

$$N = \frac{1 \text{ kg}}{3.904 \times 10^{-25} \text{ kg/atom}} = \boxed{2.56 \times 10^{24} \text{ atoms}}$$

b) The energy released in the fissioning of 1 kg of ^{235}U is

$$E = NQ = (2.56 \times 10^{24}) (2 \times 10^8 \text{ eV}) \left(\frac{1.6 \times 10^{-19} \text{ J}}{1 \text{ eV}} \right) \\ = \boxed{8.19 \times 10^{13} \text{ J}}$$

$$c) t = \frac{E}{\text{power}} = \frac{8.19 \times 10^{13} \text{ J}}{100 \text{ J/sec}} = 8.19 \times 10^{11} \text{ sec}$$

$$\text{or } T = (8.19 \times 10^{11} \text{ sec}) \left(\frac{1 \text{ yr}}{3.16 \times 10^7 \text{ sec}} \right) \\ = \boxed{2.59 \times 10^4 \text{ years}}$$

$$(44-9) Q = [(\text{mass in}) - (\text{mass out})] c^2$$

$$= [m(^{235}\text{U}) + m_n - m(^{141}\text{Cs}) - m(^{93}\text{Rb}) - 2m_n] c^2 \\ = [235.04392 \text{ u} - 1.00867 \text{ u} - 140.91963 \text{ u} - 92.92157 \text{ u}] c^2 \\ = 0.194 \text{ u} c^2$$

$$= (0.194 \times 931.5 \text{ MeV}) = \boxed{181 \text{ MeV}}$$

(44-10) This is like (44-8). The number of ^2H atoms in 1 kg is

$$N = \frac{1 \text{ kg}}{\text{mass of } ^2\text{H atom}}$$

$$= \frac{1 \text{ kg}}{(2.014102 \text{ u}) (1.661 \times 10^{-27} \text{ kg/u})} = 2.99 \times 10^{26} \text{ atoms}$$

A fusion event involves a pair of ^2H ($^2\text{H} + ^2\text{H} \rightarrow ^3\text{He} + n$), so the number of pairs of ^2H atoms is $N/2$.

Each fusion event releases $Q = 3.27 \text{ MeV}$, so the energy released is

$$\begin{aligned} E &= \left(\frac{N}{2}\right) Q \\ &= \left(\frac{2.99 \times 10^{26}}{2}\right) (3.27 \times 10^6 \text{ eV}) \left(\frac{1.6 \times 10^{-19} \text{ J}}{1 \text{ eV}}\right) \\ &= 7.82 \times 10^{13} \text{ J} \end{aligned}$$

The 100 watt lamp would burn for a time

$$t = \frac{7.82 \times 10^{13} \text{ J}}{100 \text{ J/sec}} \left(\frac{1 \text{ yr}}{3.16 \times 10^7 \text{ sec}}\right) = \boxed{2.48 \times 10^4 \text{ years}}$$

This is about 3×10^4 years.

$$(44-11) \quad Q = [(mass \text{ in}) - (mass \text{ out})] c^2$$

$$= [3m(^4\text{He}) - m(^{12}\text{C})] c^2$$

$$= [3(4.0026 \text{ u}) - 12.0000 \text{ u}] c^2$$

$$= 7.8 \times 10^{-3} \text{ u} c^2 = 7.8 \times 10^{-3} (931.5 \text{ MeV}) = \boxed{7.27 \text{ MeV}}$$

$$(44-12) \text{ a) } \frac{mass}{sec} = \frac{energy/sec}{c^2} = \frac{3.9 \times 10^{26} \text{ J/sec}}{(3 \times 10^8 \text{ m/sec})^2} = \boxed{4.33 \times 10^9 \frac{\text{kg}}{\text{sec}}}$$

b) In 4.5×10^9 yrs, the Sun has lost a mass

$$m = (4.5 \times 10^9 \text{ yrs}) \left(\frac{3.16 \times 10^7 \text{ sec}}{1 \text{ yr}}\right) (4.33 \times 10^9 \frac{\text{kg}}{\text{sec}})$$

$$= \boxed{6.16 \times 10^{26} \text{ kg}}$$

$$\text{This is only } \frac{6.16 \times 10^{26} \text{ kg}}{2.0 \times 10^{30} \text{ kg}} = \boxed{3.08 \times 10^{-4}} \text{ of the Sun's mass}$$